

Advances in the understanding of ultrarelativistic beam-plasma instabilities

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Relativistic beam-plasma instabilities are ubiquitous in high-energy astrophysics where they are thought to dissipate into heat or radiation the kinetic energy of relativistic outflows from powerful sources [1]. This occurs through the generation of intense electromagnetic fluctuations that can scatter and decelerate particles to the point of forming collisionless shocks waves or triggering bright synchrotron-type emissions. Two main instability classes are known to prevail in relativistic beam-plasma systems: the essentially electrostatic, oblique two-stream modes (OTSI) and the essentially magnetic, transverse current filamentation (CFI) modes [2]. These instabilities also operate, usually detrimentally, in laboratory schemes involving the interpenetration of relativistic electron streams and plasmas, as in plasma wakefield accelerators [3] or ultraintense laser-solid interactions [4].

This talk will cover a series of recent theoretical and numerical results on relativistic beam-plasma instabilities. Most of this work has been conducted within the E 305 project at the SLAC/FACET-II accelerator [5], which aims to probe the beam-plasma instabilities excited by an ultrarelativistic (10 GeV), high-density ($\sim 10^{19-20} \text{ cm}^{-3}$) electron beam through various gas or solid materials. First, I will report on a spatiotemporal theory, validated by particle-in-cell (PIC) simulations, describing the growth of the OTSI in collisionless plasmas permeated by finite-length beams [6]. Second, I will present a simulation study of the propagation of the FACET-II beam in a solid-density target, revealing how the hierarchy between the OTSI and CFI depends on the target collisionality and the beam density. The PIC results will be interpreted in light of a general kinetic linear theory of the OTSI and CFI modes in the presence of collisional background electrons. Finally, with a view to better understanding the physics of the precursor of relativistic collisionless shocks, I will discuss recent simulation findings on the saturation mechanisms of the CFI arising in asymmetric plasma flows [7].

References

- [1] A. M. Bykov and R. A. Treumann, *Astron. Astrophys. Rev.* **19**, 42 (2011).
- [2] A. Bret, L. Gremillet, and M. Dieckmann, *Phys. Plasmas* **17**, 120501 (2010).
- [3] C. M. Huntington *et al.*, *Phys. Rev. Lett.* **106**, 105001 (2011).
- [4] G. Raj *et al.*, *Phys. Rev. Res.* **2**, 023123 (2020).

- [5] V. Yakimenko *et al.*, Phys. Rev. Accel. Beams **22**, 101301 (2019).
- [6] P. San Miguel Claveria *et al.*, *Spatiotemporal dynamics of ultrarelativistic beam-plasma instabilities*, to be published in Phys. Rev. Res. (2022).
- [7] V. Bresci, L. Gremillet and M. Lemoine, Phys. Rev. E **105**, 035202 (2022).